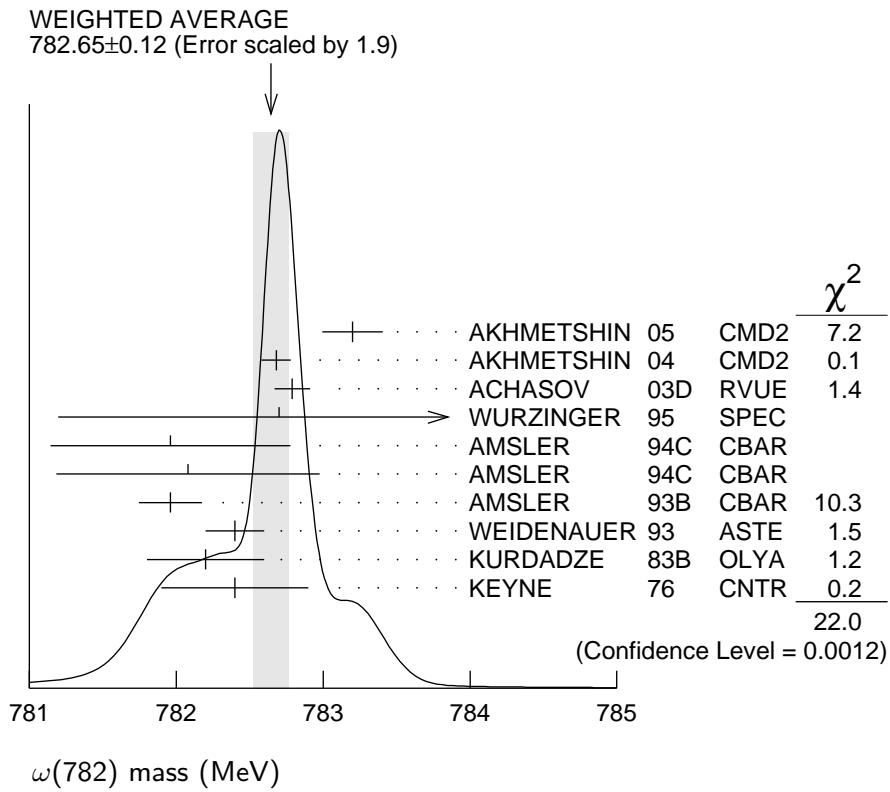


**$\omega(782)$**  $I^G(J^{PC}) = 0^-(1^{--})$  **$\omega(782)$  MASS**

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>782.65±0.12 OUR AVERAGE</b>				Error includes scale factor of 1.9. See the ideogram below.
783.20±0.13±0.16	18680	AKHMETSHIN 05	CMD2	0.60–1.38 $e^+e^- \rightarrow \pi^0\gamma$
782.68±0.09±0.04	11200	1 AKHMETSHIN 04	CMD2	$e^+e^- \rightarrow \pi^+\pi^-\pi^0$
782.79±0.08±0.09	1.2M	2 ACHASOV 03D	RVUE	$0.44–2.00 e^+e^- \rightarrow \pi^+\pi^-\pi^0$
782.7 ± 0.1 ± 1.5	19500	WURZINGER 95	SPEC	1.33 $p\bar{d} \rightarrow {}^3He\omega$
781.96±0.17±0.80	11k	3 AMSLER 94C	CBAR	0.0 $\bar{p}p \rightarrow \omega\eta\pi^0$
782.08±0.36±0.82	3463	4 AMSLER 94C	CBAR	0.0 $\bar{p}p \rightarrow \omega\eta\pi^0$
781.96±0.13±0.17	15k	AMSLER 93B	CBAR	0.0 $\bar{p}p \rightarrow \omega\pi^0\pi^0$
782.4 ± 0.2	270k	WEIDENAUER 93	ASTE	$\bar{p}p \rightarrow 2\pi^+2\pi^-\pi^0$
782.2 ± 0.4	1488	KURDADZE 83B	OLYA	$e^+e^- \rightarrow \pi^+\pi^-\pi^0$
782.4 ± 0.5	7000	5 KEYNE 76	CNTR	$\pi^-p \rightarrow \omega n$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
781.91±0.24		6 LEES 12G	BABR	$e^+e^- \rightarrow \pi^+\pi^-\gamma$
781.78±0.10		7 BARKOV 87	CMD	$e^+e^- \rightarrow \pi^+\pi^-\pi^0$
783.3 ± 0.4	433	CORDIER 80	DM1	$e^+e^- \rightarrow \pi^+\pi^-\pi^0$
782.5 ± 0.8	33260	ROOS 80	RVUE	0.0–3.6 $\bar{p}p$
782.6 ± 0.8	3000	BENKHEIRI 79	OMEG	9–12 $\pi^\pm p$
781.8 ± 0.6	1430	COOPER 78B	HBC	0.7–0.8 $\bar{p}p \rightarrow 5\pi$
782.7 ± 0.9	535	VANAPEL...	HBC	7.2 $\bar{p}p \rightarrow \bar{p}\rho\omega$
783.5 ± 0.8	2100	GESSAROLI 77	HBC	11 $\pi^-p \rightarrow \omega n$
782.5 ± 0.8	418	AGUILAR-...	HBC	3.9,4.6 $K^-p$
783.4 ± 1.0	248	BIZZARRI 71	HBC	0.0 $p\bar{p} \rightarrow K^+K^-\omega$
781.0 ± 0.6	510	BIZZARRI 71	HBC	0.0 $p\bar{p} \rightarrow K_1K_1\omega$
783.7 ± 1.0	3583	8 COYNE 71	HBC	3.7 $\pi^+p \rightarrow p\pi^+\pi^+\pi^-\pi^0$
784.1 ± 1.2	750	ABRAMOVI... 70	HBC	3.9 $\pi^-p$
783.2 ± 1.6		9 BIGGS 70B	CNTR	<4.1 $\gamma C \rightarrow \pi^+\pi^-C$
782.4 ± 0.5	2400	BIZZARRI 69	HBC	0.0 $\bar{p}p$

<sup>1</sup> Update of AKHMETSHIN 00C.<sup>2</sup> From the combined fit of ANTONELLI 92, ACHASOV 01E, ACHASOV 02E, and ACHASOV 03D data on the  $\pi^+\pi^-\pi^0$  and ANTONELLI 92 on the  $\omega\pi^+\pi^-$  final states. Supersedes ACHASOV 99E and ACHASOV 02E.<sup>3</sup> From the  $\eta \rightarrow \gamma\gamma$  decay.<sup>4</sup> From the  $\eta \rightarrow 3\pi^0$  decay.<sup>5</sup> Observed by threshold-crossing technique. Mass resolution = 4.8 MeV FWHM.<sup>6</sup> From the  $\rho - \omega$  interference in the  $\pi^+\pi^-$  mass spectrum using the Breit-Wigner for the  $\omega$  and leaving its mass and width as free parameters of the fit.<sup>7</sup> Systematic uncertainties underestimated.<sup>8</sup> From best-resolution sample of COYNE 71.<sup>9</sup> From  $\omega\rho$  interference in the  $\pi^+\pi^-$  mass spectrum assuming  $\omega$  width 12.6 MeV.



### omega(782) WIDTH

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>8.49±0.08 OUR AVERAGE</b>				
8.68±0.23±0.10	11200	<sup>1</sup> AKHMETSHIN 04	CMD2	$e^+ e^- \rightarrow \pi^+ \pi^- \pi^0$
8.68±0.04±0.15	1.2M	<sup>2</sup> ACHASOV 03D	RVUE	$0.44^{+2.00}_{-} e^+ e^- \rightarrow \pi^+ \pi^- \pi^0$
8.2 ± 0.3	19500	WURZINGER 95	SPEC	$1.33 p d \rightarrow {}^3\text{He}\omega$
8.4 ± 0.1		<sup>3</sup> AULCHENKO 87	ND	$e^+ e^- \rightarrow \pi^+ \pi^- \pi^0$
8.30±0.40		BARKOV 87	CMD	$e^+ e^- \rightarrow \pi^+ \pi^- \pi^0$
9.8 ± 0.9	1488	KURDADZE 83B	OLYA	$e^+ e^- \rightarrow \pi^+ \pi^- \pi^0$
9.0 ± 0.8	433	CORDIER 80	DM1	$e^+ e^- \rightarrow \pi^+ \pi^- \pi^0$
9.1 ± 0.8	451	BENAKSAS 72B	OSPK	$e^+ e^- \rightarrow \pi^+ \pi^- \pi^0$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
8.13±0.45		<sup>4</sup> LEES 12G	BABR	$e^+ e^- \rightarrow \pi^+ \pi^- \gamma$
12 ± 2	1430	COOPER 78B	HBC	$0.7-0.8 \bar{p}p \rightarrow 5\pi$
9.4 ± 2.5	2100	GESSAROLI 77	HBC	$11 \pi^- p \rightarrow \omega n$
10.22±0.43	20000	<sup>5</sup> KEYNE 76	CNTR	$\pi^- p \rightarrow \omega n$
13.3 ± 2	418	AGUILAR-...	HBC	$3.9, 4.6 K^- p$
10.5 ± 1.5		BORENSTEIN 72	HBC	$2.18 K^- p$
7.70±0.9 ± 1.15	940	BROWN 72	MMS	$2.5 \pi^- p \rightarrow n\text{MM}$
10.3 ± 1.4	510	BIZZARRI 71	HBC	$0.0 p\bar{p} \rightarrow K_1 K_1 \omega$
12.8 ± 3.0	248	BIZZARRI 71	HBC	$0.0 p\bar{p} \rightarrow K^+ K^- \omega$
9.5 ± 1.0	3583	COYNE 71	HBC	$3.7 \pi^+ p \rightarrow p\pi^+ \pi^+ \pi^- \pi^0$

<sup>1</sup> Update of AKHMETSHIN 00C.<sup>2</sup> From the combined fit of ANTONELLI 92, ACHASOV 01E, ACHASOV 02E, and ACHASOV 03D data on the  $\pi^+ \pi^- \pi^0$  and ANTONELLI 92 on the  $\omega \pi^+ \pi^-$  final states. Supersedes ACHASOV 99E and ACHASOV 02E.<sup>3</sup> Relativistic Breit-Wigner includes radiative corrections.<sup>4</sup> From the  $\rho - \omega$  interference in the  $\pi^+ \pi^-$  mass spectrum using the Breit-Wigner for the  $\omega$  and leaving its mass and width as free parameters of the fit.<sup>5</sup> Observed by threshold-crossing technique. Mass resolution = 4.8 MeV FWHM.

## $\omega(782)$ DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )	Scale factor/ Confidence level
$\Gamma_1 \pi^+ \pi^- \pi^0$	(89.2 $\pm$ 0.7) %	
$\Gamma_2 \pi^0 \gamma$	( 8.28 $\pm$ 0.28) %	S=2.1
$\Gamma_3 \pi^+ \pi^-$	( 1.53 $\pm$ 0.11) %	S=1.2
$\Gamma_4$ neutrals (excluding $\pi^0 \gamma$ )	( 8 $\pm$ 8 ) $\times 10^{-3}$	S=1.1
$\Gamma_5 \eta \gamma$	( 4.6 $\pm$ 0.4) $\times 10^{-4}$	S=1.1
$\Gamma_6 \pi^0 e^+ e^-$	( 7.7 $\pm$ 0.6) $\times 10^{-4}$	
$\Gamma_7 \pi^0 \mu^+ \mu^-$	( 1.3 $\pm$ 0.4) $\times 10^{-4}$	S=2.1
$\Gamma_8 \eta e^+ e^-$		
$\Gamma_9 e^+ e^-$	( 7.28 $\pm$ 0.14) $\times 10^{-5}$	S=1.3
$\Gamma_{10} \pi^+ \pi^- \pi^0 \pi^0$	< 2 $\times 10^{-4}$	CL=90%
$\Gamma_{11} \pi^+ \pi^- \gamma$	< 3.6 $\times 10^{-3}$	CL=95%
$\Gamma_{12} \pi^+ \pi^- \pi^+ \pi^-$	< 1 $\times 10^{-3}$	CL=90%
$\Gamma_{13} \pi^0 \pi^0 \gamma$	( 6.6 $\pm$ 1.1) $\times 10^{-5}$	
$\Gamma_{14} \eta \pi^0 \gamma$	< 3.3 $\times 10^{-5}$	CL=90%
$\Gamma_{15} \mu^+ \mu^-$	( 9.0 $\pm$ 3.1) $\times 10^{-5}$	
$\Gamma_{16} 3\gamma$	< 1.9 $\times 10^{-4}$	CL=95%

## Charge conjugation ( $C$ ) violating modes

$\Gamma_{17} \eta \pi^0$	$C < 2.1 \times 10^{-4}$	CL=90%
$\Gamma_{18} 2\pi^0$	$C < 2.1 \times 10^{-4}$	CL=90%
$\Gamma_{19} 3\pi^0$	$C < 2.3 \times 10^{-4}$	CL=90%

## CONSTRAINED FIT INFORMATION

An overall fit to 15 branching ratios uses 51 measurements and one constraint to determine 10 parameters. The overall fit has a  $\chi^2 = 51.8$  for 42 degrees of freedom.

The following *off-diagonal* array elements are the correlation coefficients  $\langle \delta x_i \delta x_j \rangle / (\delta x_i \cdot \delta x_j)$ , in percent, from the fit to the branching fractions,  $x_i \equiv \Gamma_i / \Gamma_{\text{total}}$ . The fit constrains the  $x_i$  whose labels appear in this array to sum to one.

$x_2$	22								
$x_3$	-18	-4							
$x_4$	-92	-56	1						
$x_5$	7	7	-1	-9					
$x_6$	-1	0	0	0	0				
$x_7$	-1	0	0	0	0	0			
$x_9$	-38	-33	7	44	-21	0	0		
$x_{13}$	1	4	0	-2	0	0	0	-1	
$x_{15}$	0	0	0	0	0	0	0	0	0
	$x_1$	$x_2$	$x_3$	$x_4$	$x_5$	$x_6$	$x_7$	$x_9$	$x_{13}$

### $\omega(782)$ PARTIAL WIDTHS

#### $\Gamma(\pi^0 \gamma)$

VALUE (keV)	EVTS	DOCUMENT ID	TECN	COMMENT	$\Gamma_2$
<b>• • •</b> We do not use the following data for averages, fits, limits, etc. <b>• • •</b>					
880 $\pm$ 50	7815	<sup>1</sup> ACHASOV	13	SND    1.05–2.00 $e^+ e^- \rightarrow \pi^0 \pi^0 \gamma$	
788 $\pm$ 12 $\pm$ 27	36500	<sup>2</sup> ACHASOV	03	SND    0.60–0.97 $e^+ e^- \rightarrow \pi^0 \gamma$	
764 $\pm$ 51	10625	DOLINSKY	89	ND $e^+ e^- \rightarrow \pi^0 \gamma$	

<sup>1</sup> Systematic uncertainty not estimated.

<sup>2</sup> Using  $\Gamma_\omega = 8.44 \pm 0.09$  MeV and  $B(\omega \rightarrow \pi^0 \gamma)$  from ACHASOV 03.

#### $\Gamma(\eta \gamma)$

VALUE (keV)	EVTS	DOCUMENT ID	TECN	COMMENT	$\Gamma_5$
<b>• • •</b> We do not use the following data for averages, fits, limits, etc. <b>• • •</b>					
6.1 $\pm$ 2.5		<sup>1</sup> DOLINSKY	89	ND $e^+ e^- \rightarrow \eta \gamma$	

<sup>1</sup> Using  $\Gamma_\omega = 8.4 \pm 0.1$  MeV and  $B(\omega \rightarrow \eta \gamma)$  from DOLINSKY 89.

#### $\Gamma(e^+ e^-)$

VALUE (keV)	EVTS	DOCUMENT ID	TECN	COMMENT	$\Gamma_9$
<b>0.60 <math>\pm</math> 0.02 OUR EVALUATION</b>					
<b>• • •</b> We do not use the following data for averages, fits, limits, etc. <b>• • •</b>					
0.591 $\pm$ 0.015	11200	<sup>1,2</sup> AKHMETSHIN 04	CMD2	$e^+ e^- \rightarrow \pi^+ \pi^- \pi^0$	
0.653 $\pm$ 0.003 $\pm$ 0.021	1.2M	<sup>3</sup> ACHASOV	03D	RVUE $0.44–2.00 \pi^+ \pi^- \pi^0$	
0.600 $\pm$ 0.031	10625	DOLINSKY	89	ND $e^+ e^- \rightarrow \pi^0 \gamma$	

<sup>1</sup> Using  $B(\omega \rightarrow \pi^+ \pi^- \pi^0) = 0.891 \pm 0.007$  and  $\Gamma_{\text{total}} = 8.44 \pm 0.09$  MeV.

<sup>2</sup> Update of AKHMETSHIN 00C.

<sup>3</sup> Using ACHASOV 03, ACHASOV 03D and  $B(\omega \rightarrow \pi^+ \pi^-) = (1.70 \pm 0.28)\%$ .

### $\omega(782) \Gamma(e^+ e^-) \Gamma(i)/\Gamma^2(\text{total})$

$\Gamma(e^+ e^-)/\Gamma_{\text{total}} \times \Gamma(\pi^+ \pi^- \pi^0)/\Gamma_{\text{total}}$	$\Gamma_9/\Gamma \times \Gamma_1/\Gamma$
<i>VALUE (units <math>10^{-5}</math>)</i>	<i>EVTS</i> <i>DOCUMENT ID</i> <i>TECN</i> <i>COMMENT</i>

**6.49±0.11 OUR FIT** Error includes scale factor of 1.3.

**6.38±0.10 OUR AVERAGE** Error includes scale factor of 1.1.

6.24±0.11±0.08	11.2k	<sup>1</sup> AKHMETSHIN 04	CMD2	$e^+ e^- \rightarrow \pi^+ \pi^- \pi^0$
6.70±0.06±0.27		AUBERT,B	04N	$BABR \quad 10.6 \quad e^+ e^- \rightarrow \pi^+ \pi^- \pi^0 \gamma$
6.74±0.04±0.24	1.2M	<sup>2,3</sup> ACHASOV	03D	$RVUE \quad 0.44-2.00 \quad e^+ e^- \rightarrow \pi^+ \pi^- \pi^0$
6.37±0.35		<sup>2</sup> DOLINSKY	89	$ND \quad e^+ e^- \rightarrow \pi^+ \pi^- \pi^0$
6.45±0.24		<sup>2</sup> BARKOV	87	$CMD \quad e^+ e^- \rightarrow \pi^+ \pi^- \pi^0$
5.79±0.42	1488	<sup>2</sup> KURDADZE	83B	$OLYA \quad e^+ e^- \rightarrow \pi^+ \pi^- \pi^0$
5.89±0.54	433	<sup>2</sup> CORDIER	80	$DM1 \quad e^+ e^- \rightarrow \pi^+ \pi^- \pi^0$
7.54±0.84	451	<sup>2</sup> BENAKSAS	72B	$OSPK \quad e^+ e^- \rightarrow \pi^+ \pi^- \pi^0$

• • • We do not use the following data for averages, fits, limits, etc. • • •

6.20±0.13		<sup>4</sup> BENAYOUN	10	$RVUE \quad 0.4-1.05 \quad e^+ e^-$
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<sup>1</sup> Update of AKHMETSHIN 00C.

<sup>2</sup> Recalculated by us from the cross section in the peak.

<sup>3</sup> From the combined fit of ANTONELLI 92, ACHASOV 01E, ACHASOV 02E, and ACHASOV 03D data on the  $\pi^+ \pi^- \pi^0$  and ANTONELLI 92 on the  $\omega \pi^+ \pi^-$  final states. Supersedes ACHASOV 99E and ACHASOV 02E.

<sup>4</sup> A simultaneous fit of  $e^+ e^- \rightarrow \pi^+ \pi^-, \pi^+ \pi^- \pi^0, \pi^0 \gamma, \eta \gamma$  data.

$\Gamma(e^+ e^-)/\Gamma_{\text{total}} \times \Gamma(\pi^0 \gamma)/\Gamma_{\text{total}}$	$\Gamma_9/\Gamma \times \Gamma_2/\Gamma$
<i>VALUE (units <math>10^{-6}</math>)</i>	<i>EVTS</i> <i>DOCUMENT ID</i> <i>TECN</i> <i>COMMENT</i>

**6.02±0.20 OUR FIT** Error includes scale factor of 1.9.

**6.45±0.17 OUR AVERAGE**

6.47±0.14±0.39	18680	AKHMETSHIN 05	CMD2	$0.60-1.38 \quad e^+ e^- \rightarrow \pi^0 \gamma$
6.50±0.11±0.20	36500	<sup>1</sup> ACHASOV	03	$SND \quad 0.60-0.97 \quad e^+ e^- \rightarrow \pi^0 \gamma$
6.34±0.21±0.21	10625	<sup>2</sup> DOLINSKY	89	$ND \quad e^+ e^- \rightarrow \pi^0 \gamma$

• • • We do not use the following data for averages, fits, limits, etc. • • •

6.80±0.13		<sup>3</sup> BENAYOUN	10	$RVUE \quad 0.4-1.05 \quad e^+ e^-$
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<sup>1</sup> Using  $\sigma_{\phi \rightarrow \pi^0 \gamma}$  from ACHASOV 00 and  $m_\omega = 782.57$  MeV in the model with the energy-independent phase of  $\rho$ - $\omega$  interference equal to  $(-10.2 \pm 7.0)^\circ$ .

<sup>2</sup> Recalculated by us from the cross section in the peak.

<sup>3</sup> A simultaneous fit of  $e^+ e^- \rightarrow \pi^+ \pi^-, \pi^+ \pi^- \pi^0, \pi^0 \gamma, \eta \gamma$  data.

$\Gamma(e^+e^-)/\Gamma_{\text{total}} \times \Gamma(\pi^+\pi^-)/\Gamma_{\text{total}}$	$\Gamma_9/\Gamma \times \Gamma_3/\Gamma$			
<u>VALUE (units <math>10^{-6}</math>)</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>1.225±0.058±0.041</b>	800k	1 ACHASOV 06	SND	$e^+e^- \rightarrow \pi^+\pi^-$
<b>• • • We do not use the following data for averages, fits, limits, etc. • • •</b>				
1.166±0.036		2 BENAYOUN 13	RVUE	$0.4-1.05 e^+e^-$
1.05 ± 0.08		3 DAVIER 13	RVUE	$e^+e^- \rightarrow \pi^+\pi^-(\gamma)$
<b>1 Supersedes ACHASOV 05A.</b>				
<b>2 A simultaneous fit to <math>e^+e^- \rightarrow \pi^+\pi^-, \pi^+\pi^-\pi^0, \pi^0\gamma, \eta\gamma, K\bar{K}</math>, and <math>\tau^- \rightarrow \pi^-\pi^0\nu_\tau</math> data. Supersedes BENAYOUN 10.</b>				
<b>3 From <math>e^+e^- \rightarrow \pi^+\pi^-(\gamma)</math> data of LEES 12G.</b>				

$\Gamma(e^+e^-)/\Gamma_{\text{total}} \times \Gamma(\eta\gamma)/\Gamma_{\text{total}}$	$\Gamma_9/\Gamma \times \Gamma_5/\Gamma$			
<u>VALUE (units <math>10^{-8}</math>)</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>3.32±0.28 OUR FIT</b>	Error includes scale factor of 1.1.			
<b>3.18±0.28 OUR AVERAGE</b>				
3.10±0.31±0.11	33k	1 ACHASOV 07B	SND	$0.6-1.38 e^+e^- \rightarrow \eta\gamma$
3.17 <sup>+1.85</sup> <sub>-1.31</sub> ± 0.21	17.4k	2 AKHMETSHIN 05	CMD2	$0.60-1.38 e^+e^- \rightarrow \eta\gamma$
3.41±0.52±0.21	23k	3,4 AKHMETSHIN 01B	CMD2	$e^+e^- \rightarrow \eta\gamma$
<b>• • • We do not use the following data for averages, fits, limits, etc. • • •</b>				
4.50±0.10		5 BENAYOUN 10	RVUE	$0.4-1.05 e^+e^-$
<b>1 From a combined fit of <math>\sigma(e^+e^- \rightarrow \eta\gamma)</math> with <math>\eta \rightarrow 3\pi^0</math> and <math>\eta \rightarrow \pi^+\pi^-\pi^0</math>, and fixing <math>B(\eta \rightarrow 3\pi^0) / B(\eta \rightarrow \pi^+\pi^-\pi^0) = 1.44 \pm 0.04</math>. Recalculated by us from the cross section at the peak. Supersedes ACHASOV 00D and ACHASOV 06A.</b>				
<b>2 From the <math>\eta \rightarrow 2\gamma</math> decay and using <math>B(\eta \rightarrow \gamma\gamma) = 39.43 \pm 0.26\%</math>.</b>				
<b>3 From the <math>\eta \rightarrow 3\pi^0</math> decay and using <math>B(\eta \rightarrow 3\pi^0) = (32.24 \pm 0.29) \times 10^{-2}</math>.</b>				
<b>4 The combined fit from 600 to 1380 MeV taking into account <math>\rho(770)</math>, <math>\omega(782)</math>, <math>\phi(1020)</math>, and <math>\rho(1450)</math> (mass and width fixed at 1450 MeV and 310 MeV respectively).</b>				
<b>5 A simultaneous fit of <math>e^+e^- \rightarrow \pi^+\pi^-, \pi^+\pi^-\pi^0, \pi^0\gamma, \eta\gamma</math> data.</b>				

## $\omega(782)$ BRANCHING RATIOS

$\Gamma(\pi^+\pi^-\pi^0)/\Gamma_{\text{total}}$	$\Gamma_1/\Gamma$			
<b>NIECKNIG 12 describes final-state interactions between the three pions in a dispersive framework using data on the <math>\pi\pi</math> P-wave scattering phase shift.</b>				
<u>VALUE</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>• • • We do not use the following data for averages, fits, limits, etc. • • •</b>				
0.9024±0.0019		1 AMBROSINO 08G	KLOE	$1.0-1.03 e^+e^- \rightarrow \pi^+\pi^-\pi^0, 2\pi^0\gamma$
0.8965±0.0016±0.0048	1.2M	2,3 ACHASOV 03D	RVUE	$0.44-2.00 e^+e^- \rightarrow \pi^+\pi^-\pi^0$
0.880 ± 0.020 ± 0.032	11200	3,4 AKHMETSHIN 00C	CMD2	$e^+e^- \rightarrow \pi^+\pi^-\pi^0$
0.8942±0.0062		3 DOLINSKY 89	ND	$e^+e^- \rightarrow \pi^+\pi^-\pi^0$
<b>1 Not independent of <math>\Gamma(\pi^0\gamma) / \Gamma(\pi^+\pi^-\pi^0)</math> from AMBROSINO 08G.</b>				
<b>2 Using ACHASOV 03, ACHASOV 03D and <math>B(\omega \rightarrow \pi^+\pi^-) = (1.70 \pm 0.28)\%</math>.</b>				
<b>3 Not independent of the corresponding <math>\Gamma(e^+e^-) \times \Gamma(\pi^+\pi^-\pi^0)/\Gamma_{\text{total}}^2</math>.</b>				
<b>4 Using <math>\Gamma(e^+e^-) = 0.60 \pm 0.02</math> keV.</b>				

$\Gamma(\pi^0\gamma)/\Gamma_{\text{total}}$  $\Gamma_2/\Gamma$ 

<u>VALUE (units <math>10^{-2}</math>)</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>• • • We do not use the following data for averages, fits, limits, etc. • • •</b>				
8.09 $\pm 0.14$		<sup>1</sup> AMBROSINO 08G	KLOE	$e^+e^- \rightarrow \pi^+\pi^-2\pi^0, 2\pi^0\gamma$
9.06 $\pm 0.20 \pm 0.57$	18680	<sup>2,3</sup> AKHMETSHIN 05	CMD2	$0.60-1.38 e^+e^- \rightarrow \pi^0\gamma$
9.34 $\pm 0.15 \pm 0.31$	36500	<sup>3</sup> ACHASOV 03	SND	$0.60-0.97 e^+e^- \rightarrow \pi^0\gamma$
8.65 $\pm 0.16 \pm 0.42$	1.2M	<sup>4,5</sup> ACHASOV 03D	RVUE	$0.44-2.00 e^+e^- \rightarrow \pi^+\pi^-\pi^0$
8.39 $\pm 0.24$	9975	<sup>6</sup> BENAYOUN 96	RVUE	$e^+e^- \rightarrow \pi^0\gamma$
8.88 $\pm 0.62$	10625	<sup>3</sup> DOLINSKY 89	ND	$e^+e^- \rightarrow \pi^0\gamma$

<sup>1</sup> Not independent of  $\Gamma(\pi^0\gamma) / \Gamma(\pi^+\pi^-\pi^0)$  from AMBROSINO 08G.<sup>2</sup> Using  $B(\omega \rightarrow e^+e^-) = (7.14 \pm 0.13) \times 10^{-5}$ .<sup>3</sup> Not independent of the corresponding  $\Gamma(e^+e^-) \times \Gamma(\pi^0\gamma)/\Gamma_{\text{total}}^2$ .<sup>4</sup> Using ACHASOV 03, ACHASOV 03D and  $B(\omega \rightarrow \pi^+\pi^-) = (1.70 \pm 0.28)\%$ .<sup>5</sup> Not independent of the corresponding  $\Gamma(e^+e^-) \times \Gamma(\pi^+\pi^-\pi^0)/\Gamma_{\text{total}}^2$ .<sup>6</sup> Reanalysis of DRUZHININ 84, DOLINSKY 89, DOLINSKY 91 taking into account the triangle anomaly contributions. $\Gamma(\pi^0\gamma)/\Gamma(\pi^+\pi^-\pi^0)$  $\Gamma_2/\Gamma_1$ 

<u>VALUE (units <math>10^{-2}</math>)</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>9.28 <math>\pm 0.31</math> OUR FIT</b>	Error includes scale factor of 2.3.		
<b>9.05 <math>\pm 0.27</math> OUR AVERAGE</b>	Error includes scale factor of 1.8.		
8.97 $\pm 0.16$	AMBROSINO 08G	KLOE	$e^+e^- \rightarrow \pi^+\pi^-2\pi^0, 2\pi^0\gamma$
9.94 $\pm 0.36 \pm 0.38$	<sup>1</sup> AULCHENKO 00A	SND	$e^+e^- \rightarrow \pi^+\pi^-2\pi^0, 2\pi^0\gamma$
8.4 $\pm 1.3$	KEYNE 76	CNTR	$\pi^-p \rightarrow \omega n$
10.9 $\pm 2.5$	BENAKSAS 72C	OSPK	$e^+e^- \rightarrow \pi^0\gamma$
8.1 $\pm 2.0$	BALDIN 71	HLBC	$2.9 \pi^+p$
13 $\pm 4$	JACQUET 69B	HLBC	$2.05 \pi^+p \rightarrow \pi^+p\omega$
<b>• • • We do not use the following data for averages, fits, limits, etc. • • •</b>			
9.7 $\pm 0.2 \pm 0.5$	<sup>2,3</sup> ACHASOV 03D	RVUE	$0.44-2.00 e^+e^- \rightarrow \pi^+\pi^-\pi^0$
9.9 $\pm 0.7$	<sup>2</sup> DOLINSKY 89	ND	$e^+e^- \rightarrow \pi^0\gamma$

<sup>1</sup> From  $\sigma_0^{\omega\pi^0 \rightarrow \pi^0\pi^0\gamma}(m_\phi)/\sigma_0^{\omega\pi^0 \rightarrow \pi^+\pi^-\pi^0\pi^0}(m_\phi)$  with a phase-space correction factor of 1/1.023.<sup>2</sup> Not independent of the corresponding  $\Gamma(e^+e^-) \times \Gamma(\pi^0\gamma)/\Gamma_{\text{total}}^2$ .<sup>3</sup> Using ACHASOV 03. Based on 1.2M events. $\Gamma(\pi^+\pi^-)/\Gamma_{\text{total}}$  $\Gamma_3/\Gamma$ See also  $\Gamma(\pi^+\pi^-)/\Gamma(\pi^+\pi^-\pi^0)$ .

<u>VALUE (units <math>10^{-2}</math>)</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>1.53 <math>\pm 0.11</math> OUR FIT</b>	Error includes scale factor of 1.2.			
<b>1.49 <math>\pm 0.13</math> OUR AVERAGE</b>	Error includes scale factor of 1.3. See the ideogram below.			
1.46 $\pm 0.12 \pm 0.02$	900k	<sup>1</sup> AKHMETSHIN 07		$e^+e^- \rightarrow \pi^+\pi^-$
1.30 $\pm 0.24 \pm 0.05$	11.2k	<sup>2</sup> AKHMETSHIN 04	CMD2	$e^+e^- \rightarrow \pi^+\pi^-$
2.38 $\pm 1.77 \pm 0.18$	5.4k	<sup>3</sup> ACHASOV 02E	SND	$1.1-1.38 e^+e^- \rightarrow \pi^+\pi^-\pi^0$
2.3 $\pm 0.5$		BARKOV 85	OLYA	$e^+e^- \rightarrow \pi^+\pi^-$
1.6 $\pm 0.9$	-0.7	QUENZER 78	DM1	$e^+e^- \rightarrow \pi^+\pi^-$
3.6 $\pm 1.9$		BENAKSAS 72	OSPK	$e^+e^- \rightarrow \pi^+\pi^-$

• • • We do not use the following data for averages, fits, limits, etc. • • •

$1.75 \pm 0.11$	$4.5M$	<sup>4</sup> ACHASOV	05A	SND	$e^+ e^- \rightarrow \pi^+ \pi^-$
$2.01 \pm 0.29$		<sup>5</sup> BENAYOUN	03	RVUE	$e^+ e^- \rightarrow \pi^+ \pi^-$
$1.9 \pm 0.3$		<sup>6</sup> GARDNER	99	RVUE	$e^+ e^- \rightarrow \pi^+ \pi^-$
$2.3 \pm 0.4$		<sup>7</sup> BENAYOUN	98	RVUE	$e^+ e^- \rightarrow \pi^+ \pi^-, \mu^+ \mu^-$
$1.0 \pm 0.11$		<sup>8</sup> WICKLUND	78	ASPK	$3,4,6 \pi^\pm N$
$1.22 \pm 0.30$		ALVENSLEB...	71C	CNTR	Photoproduction
$1.3^{+1.2}_{-0.9}$		MOFFEIT	71	HBC	$2.8,4.7 \gamma p$
$0.80^{+0.28}_{-0.20}$		<sup>9</sup> BIGGS	70B	CNTR	$4.2 \gamma C \rightarrow \pi^+ \pi^- C$

<sup>1</sup> A combined fit of AKHMETSHIN 07, AULCHENKO 06, and AULCHENKO 05.

<sup>2</sup> Update of AKHMETSHIN 02.

<sup>3</sup> From the  $m_{\pi^+ \pi^-}$  spectrum taking into account the interference of the  $\rho\pi$  and  $\omega\pi$  amplitudes.

<sup>4</sup> Using  $\Gamma(\omega \rightarrow e^+ e^-)$  from the 2004 Edition of this Review (PDG 04).

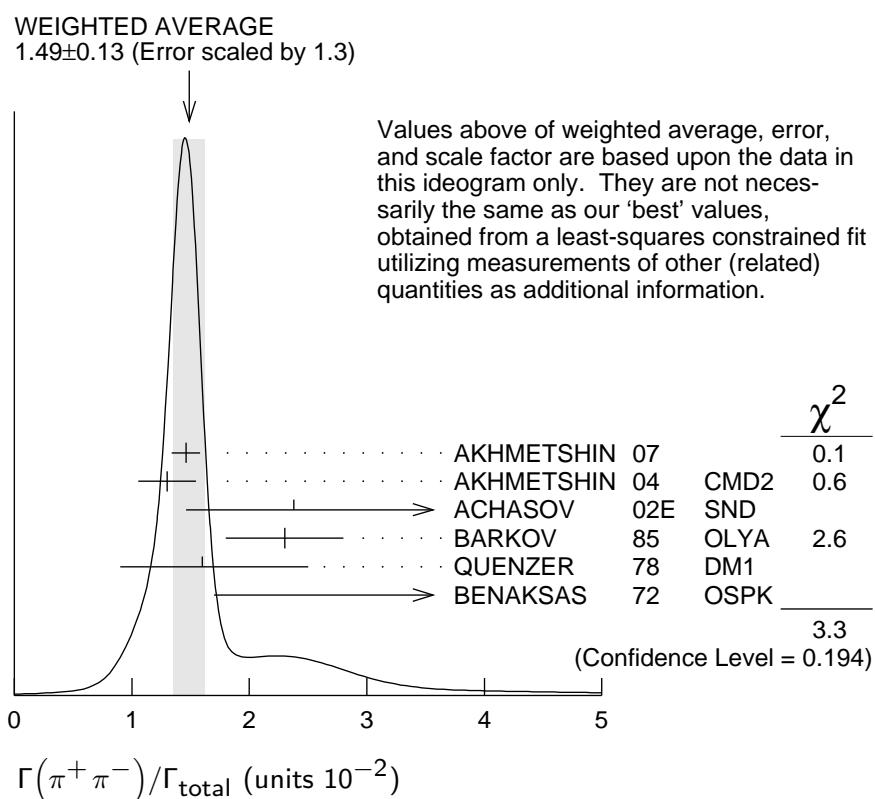
<sup>5</sup> Using the data of AKHMETSHIN 02 in the hidden local symmetry model.

<sup>6</sup> Using the data of BARKOV 85.

<sup>7</sup> Using the data of BARKOV 85 in the hidden local symmetry model.

<sup>8</sup> From a model-dependent analysis assuming complete coherence.

<sup>9</sup> Re-evaluated under  $\Gamma(\pi^+ \pi^-)/\Gamma(\pi^+ \pi^- \pi^0)$  by BEHREND 71 using more accurate  $\omega \rightarrow \rho$  photoproduction cross-section ratio.



$\Gamma(\pi^+\pi^-)/\Gamma(\pi^+\pi^-\pi^0)$  $\Gamma_3/\Gamma_1$ See also  $\Gamma(\pi^+\pi^-)/\Gamma_{\text{total}}$ .

<u>VALUE</u>		<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>0.0172±0.0014 OUR FIT</b>	Error includes scale factor of 1.2.			
<b>0.026 ±0.005 OUR AVERAGE</b>				
0.021 +0.028 -0.009	1,2 RATCLIFF	72	ASPK	$15 \pi^- p \rightarrow n 2\pi$
0.028 ±0.006	1 BEHREND	71	ASPK	Photoproduction
0.022 +0.009 -0.01	3 ROOS	70	RVUE	

<sup>1</sup> The fitted width of these data is 160 MeV in agreement with present average, thus the  $\omega$  contribution is overestimated. Assuming  $\rho$  width 145 MeV.

<sup>2</sup> Significant interference effect observed. NB of  $\omega \rightarrow 3\pi$  comes from an extrapolation.

<sup>3</sup> ROOS 70 combines ABRAMOVICH 70 and BIZZARRI 70.

 $\Gamma(\pi^+\pi^-)/\Gamma(\pi^0\gamma)$  $\Gamma_3/\Gamma_2$ 

<u>VALUE</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>0.20±0.04</b>	1.98M	1 ALOISIO	03	KLOE $1.02 e^+ e^- \rightarrow \pi^+ \pi^- \pi^0$

<sup>1</sup> Using the data of ALOISIO 02D.

 $\Gamma(\text{ neutrals})/\Gamma_{\text{total}}$  $(\Gamma_2+\Gamma_4)/\Gamma$ 

<u>VALUE</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>0.091±0.006 OUR FIT</b>				
<b>0.081±0.011 OUR AVERAGE</b>				
0.075±0.025		BIZZARRI	71	HBC $0.0 p\bar{p}$
0.079±0.019		DEINET	69B	OSPK $1.5 \pi^- p$
0.084±0.015		BOLLINI	68C	CNTR $2.1 \pi^- p$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.073±0.018	42	BASILE	72B	CNTR $1.67 \pi^- p$

 $\Gamma(\text{ neutrals})/\Gamma(\pi^+\pi^-\pi^0)$  $(\Gamma_2+\Gamma_4)/\Gamma_1$ 

<u>VALUE</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>0.102±0.008 OUR FIT</b>				
<b>0.103<sup>+0.011</sup><sub>-0.010</sub> OUR AVERAGE</b>				
0.15 ±0.04	46	AGUILAR...	72B	HBC $3.9, 4.6 K^- p$
0.10 ±0.03	19	BARASH	67B	HBC $0.0 \bar{p}p$
0.134±0.026	850	DIGIUGNO	66B	CNTR $1.4 \pi^- p$
0.097±0.016	348	FLATTE	66	HBC $1.4 - 1.7 K^- p \rightarrow \Lambda MM$
0.06 <sup>+0.05</sup> <sub>-0.02</sub>		JAMES	66	HBC $2.1 \pi^+ p$
0.08 ±0.03	35	KRAEMER	64	DBC $1.2 \pi^+ d$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.11 ±0.02	20	BUSCHBECK	63	HBC $1.5 K^- p$

### $\Gamma(\pi^0\gamma)/\Gamma(\text{ neutrals})$

### $\Gamma_2/(\Gamma_2+\Gamma_4)$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>• • • We do not use the following data for averages, fits, limits, etc. • • •</b>				
0.78±0.07		<sup>1</sup> DAKIN	72	OSPK $1.4 \pi^- p \rightarrow n\text{MM}$
>0.81	90	DEINET	69B	OSPK

<sup>1</sup> Error statistical only. Authors obtain good fit also assuming  $\pi^0\gamma$  as the only neutral decay.

### $\Gamma(\text{ neutrals})/\Gamma(\text{ charged particles})$

### $(\Gamma_2+\Gamma_4)/(\Gamma_1+\Gamma_3)$

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.100±0.008 OUR FIT</b>			
<b>0.124±0.021</b>	FELDMAN	67C	OSPK $1.2 \pi^- p$

### $\Gamma(\eta\gamma)/\Gamma_{\text{total}}$

### $\Gamma_5/\Gamma$

VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>4.6 ±0.4 OUR FIT</b>				Error includes scale factor of 1.1.
<b>6.3 ±1.3 OUR AVERAGE</b>				Error includes scale factor of 1.2.
6.6 ±1.7		<sup>1</sup> ABELE	97E	CBAR $0.0 \bar{p}p \rightarrow 5\gamma$

8.3 ±2.1 ALDE 93 GAM2  $38\pi^- p \rightarrow \omega n$

3.0  $^{+2.5}_{-1.8}$  <sup>2</sup> ANDREWS 77 CNTR 6.7–10  $\gamma\text{Cu}$

**• • • We do not use the following data for averages, fits, limits, etc. • • •**

4.3 ±0.5 ±0.1 33k <sup>3</sup> ACHASOV 07B SND 0.6–1.38  $e^+e^- \rightarrow \eta\gamma$

<sup>4,5</sup> AKHMETSHIN 05 CMD2 0.60–1.38  $e^+e^- \rightarrow \eta\gamma$

5.10±0.72±0.34 23k <sup>6</sup> AKHMETSHIN 01B CMD2  $e^+e^- \rightarrow \eta\gamma$

0.7 to 5.5 CASE 00 CBAR 0.0  $p\bar{p} \rightarrow \eta\eta\gamma$

6.56  $^{+2.41}_{-2.55}$  3525 <sup>2,8</sup> BENAYOUN 96 RVUE  $e^+e^- \rightarrow \eta\gamma$

7.3 ±2.9 <sup>2,4</sup> DOLINSKY 89 ND  $e^+e^- \rightarrow \eta\gamma$

<sup>1</sup> No flat  $\eta\eta\gamma$  background assumed.

<sup>2</sup> Solution corresponding to constructive  $\omega$ - $\rho$  interference.

<sup>3</sup> ACHASOV 07B reports  $[\Gamma(\omega(782) \rightarrow \eta\gamma)/\Gamma_{\text{total}}] \times [B(\omega(782) \rightarrow e^+e^-)] = (3.10 \pm 0.31 \pm 0.11) \times 10^{-8}$  which we divide by our best value  $B(\omega(782) \rightarrow e^+e^-) = (7.28 \pm 0.14) \times 10^{-5}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value. Supersedes ACHASOV 00D and ACHASOV 06A.

<sup>4</sup> Not independent of the corresponding  $\Gamma(e^+e^-) \times \Gamma(\eta\gamma)/\Gamma_{\text{total}}^2$ .

<sup>5</sup> Using  $B(\omega \rightarrow e^+e^-) = (7.14 \pm 0.13) \times 10^{-5}$  and  $B(\eta \rightarrow \gamma\gamma) = 39.43 \pm 0.26\%$ .

<sup>6</sup> Using  $B(\omega \rightarrow e^+e^-) = (7.07 \pm 0.19) \times 10^{-5}$  and using  $B(\eta \rightarrow 3\pi^0) = (32.24 \pm 0.29) \times 10^{-2}$ . Solution corresponding to constructive  $\omega$ - $\rho$  interference. The combined fit from 600 to 1380 MeV taking into account  $\rho(770)$ ,  $\omega(782)$ ,  $\phi(1020)$ , and  $\rho(1450)$  (mass and width fixed at 1450 MeV and 310 MeV respectively). Not independent of the corresponding  $\Gamma(e^+e^-) \times \Gamma(\eta\gamma)/\Gamma_{\text{total}}^2$ .

<sup>7</sup> Depending on the degree of coherence with the flat  $\eta\eta\gamma$  background and using  $B(\omega \rightarrow \pi^0\gamma) = (8.5 \pm 0.5) \times 10^{-2}$ .

<sup>8</sup> Reanalysis of DRUZHININ 84, DOLINSKY 89, DOLINSKY 91 taking into account the triangle anomaly contributions.

$\Gamma(\eta\gamma)/\Gamma(\pi^0\gamma)$  $\Gamma_5/\Gamma_2$ 

<u>VALUE</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>• • • We do not use the following data for averages, fits, limits, etc. • • •</b>			
0.0098±0.0024	<sup>1</sup> ALDE	93	GAM2 $38\pi^- p \rightarrow \omega n$
0.0082±0.0033	<sup>2</sup> DOLINSKY	89	ND $e^+ e^- \rightarrow \eta\gamma$
0.010 ±0.045	APEL	72B	OSPK 4–8 $\pi^- p \rightarrow n3\gamma$

<sup>1</sup> Model independent determination.<sup>2</sup> Solution corresponding to constructive  $\omega$ - $\rho$  interference. $\Gamma(\pi^0e^+e^-)/\Gamma_{\text{total}}$  $\Gamma_6/\Gamma$ 

<u>VALUE (units <math>10^{-4}</math>)</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>7.7 ±0.6 OUR FIT</b>				
<b>7.7 ±0.6 OUR AVERAGE</b>				
7.61±0.53±0.64		ACHASOV 08	SND	$0.36\text{--}0.97 e^+ e^- \rightarrow \pi^0 e^+ e^-$
8.19±0.71±0.62		AKHMETSHIN 05A	CMD2	$0.72\text{--}0.84 e^+ e^-$
5.9 ±1.9	43	DOLINSKY 88	ND	$e^+ e^- \rightarrow \pi^0 e^+ e^-$

 $\Gamma(\pi^0\mu^+\mu^-)/\Gamma_{\text{total}}$  $\Gamma_7/\Gamma$ 

<u>VALUE (units <math>10^{-4}</math>)</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>1.3 ±0.4 OUR FIT</b> Error includes scale factor of 2.1.				
<b>1.3 ±0.4 OUR AVERAGE</b> Error includes scale factor of 2.1.				
1.72±0.25±0.14	3k	ARNALDI 09	NA60	158A In–In collisions
0.96±0.23		DZHELYADIN 81B	CNTR	$25\text{--}33 \pi^- p \rightarrow \omega n$

 $\Gamma(\eta e^+e^-)/\Gamma_{\text{total}}$  $\Gamma_8/\Gamma$ 

<u>VALUE (units <math>10^{-5}</math>)</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>• • • We do not use the following data for averages, fits, limits, etc. • • •</b>			
<1.1	AKHMETSHIN 05A	CMD2	$0.72\text{--}0.84 e^+ e^-$

 $\Gamma(e^+e^-)/\Gamma_{\text{total}}$  $\Gamma_9/\Gamma$ 

<u>VALUE (units <math>10^{-4}</math>)</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>0.728±0.014 OUR FIT</b> Error includes scale factor of 1.3.				
<b>• • • We do not use the following data for averages, fits, limits, etc. • • •</b>				
0.700±0.016	11200	<sup>1,2</sup> AKHMETSHIN 04	CMD2	$e^+ e^- \rightarrow \pi^+ \pi^- \pi^0$
0.752±0.004±0.024	1.2M	<sup>2,3</sup> ACHASOV 03D	RVUE	$0.44\text{--}2.00 e^+ e^- \rightarrow \pi^+ \pi^- \pi^0$
0.714±0.036		<sup>2</sup> DOLINSKY 89	ND	$e^+ e^- \rightarrow \pi^+ \pi^- \pi^0$
0.72 ±0.03		<sup>2</sup> BARKOV 87	CMD	$e^+ e^- \rightarrow \pi^+ \pi^- \pi^0$
0.64 ±0.04	1488	<sup>2</sup> KURDADZE 83B	OLYA	$e^+ e^- \rightarrow \pi^+ \pi^- \pi^0$
0.675±0.069	433	<sup>2</sup> CORDIER 80	DM1	$e^+ e^- \rightarrow \pi^+ \pi^- \pi^0$
0.83 ±0.10	451	<sup>2</sup> BENAKSAS 72B	OSPK	$e^+ e^- \rightarrow \pi^+ \pi^- \pi^0$
0.77 ±0.06		<sup>4</sup> AUGUSTIN 69D	OSPK	$e^+ e^- \rightarrow \pi^+ \pi^- \pi^0$
0.65 ±0.13	33	<sup>5</sup> ASTVACAT... 68	OSPK	Assume SU(3)+mixing

<sup>1</sup> Using  $B(\omega \rightarrow \pi^+ \pi^- \pi^0) = 0.891 \pm 0.007$ . Update of AKHMETSHIN 00C.<sup>2</sup> Not independent of the corresponding  $\Gamma(e^+e^-) \times \Gamma(\pi^+\pi^-\pi^0)/\Gamma_{\text{total}}^2$ .<sup>3</sup> Using ACHASOV 03, ACHASOV 03D and  $B(\omega \rightarrow \pi^+ \pi^-) = (1.70 \pm 0.28)\%$ .<sup>4</sup> Rescaled by us to correspond to  $\omega$  width 8.4 MeV. Systematic errors underestimated.<sup>5</sup> Not resolved from  $\rho$  decay. Error statistical only.

$\Gamma(\pi^+\pi^-\pi^0\pi^0)/\Gamma_{\text{total}}$   $\Gamma_{10}/\Gamma$ 

<u>VALUE</u> (units $10^{-4}$ )	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
< 2	90	ACHASOV 09A	SND	$e^+e^- \rightarrow \pi^+\pi^-\pi^0\pi^0$
$\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$				
<200	90	KURDADZE 86	OLYA	$e^+e^- \rightarrow \pi^+\pi^-\pi^0\pi^0$

 $\Gamma(\pi^+\pi^-\gamma)/\Gamma_{\text{total}}$   $\Gamma_{11}/\Gamma$ 

<u>VALUE</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<0.0036	95	WEIDENAUER 90	ASTE	$p\bar{p} \rightarrow \pi^+\pi^-\pi^+\pi^-\gamma$
$\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$				
<0.004	95	BITYUKOV 88B	SPEC	$32\pi^- p \rightarrow \pi^+\pi^-\gamma X$

 $\Gamma(\pi^+\pi^-\gamma)/\Gamma(\pi^+\pi^-\pi^0)$   $\Gamma_{11}/\Gamma_1$ 

<u>VALUE</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
$\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$				
<0.066	90	KALBFLEISCH 75	HBC	$2.18K^- p \rightarrow \Lambda\pi^+\pi^-\gamma$
<0.05	90	FLATTE 66	HBC	$1.2 - 1.7K^- p \rightarrow \Lambda\pi^+\pi^-\gamma$

 $\Gamma(\pi^+\pi^-\pi^+\pi^-)/\Gamma_{\text{total}}$   $\Gamma_{12}/\Gamma$ 

<u>VALUE</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
< $1 \times 10^{-3}$	90	KURDADZE 88	OLYA	$e^+e^- \rightarrow \pi^+\pi^-\pi^+\pi^-$

 $\Gamma(\pi^0\pi^0\gamma)/\Gamma_{\text{total}}$   $\Gamma_{13}/\Gamma$ 

<u>VALUE</u> (units $10^{-5}$ )	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b><math>6.6 \pm 1.1</math> OUR FIT</b>				
<b><math>6.5 \pm 1.2</math> OUR AVERAGE</b>				

$6.4^{+2.4}_{-2.0} \pm 0.8$	190	<sup>1</sup> AKHMETSHIN 04B	CMD2	$0.6 - 0.97 e^+e^- \rightarrow \pi^0\pi^0\gamma$
$6.6^{+1.4}_{-1.3} \pm 0.6$	295	ACHASOV 02F	SND	$0.36 - 0.97 e^+e^- \rightarrow \pi^0\pi^0\gamma$

$\bullet \bullet \bullet$  We do not use the following data for averages, fits, limits, etc.  $\bullet \bullet \bullet$

$11.8^{+2.1}_{-1.9} \pm 1.4$	190	<sup>2</sup> AKHMETSHIN 04B	CMD2	$0.6 - 0.97 e^+e^- \rightarrow \pi^0\pi^0\gamma$
$7.8 \pm 2.7 \pm 2.0$	63	1,3 ACHASOV 00G	SND	$e^+e^- \rightarrow \pi^0\pi^0\gamma$
$12.7 \pm 2.3 \pm 2.5$	63	2,3 ACHASOV 00G	SND	$e^+e^- \rightarrow \pi^0\pi^0\gamma$

<sup>1</sup>In the model assuming the  $\rho \rightarrow \pi^0\pi^0\gamma$  decay via the  $\omega\pi$  and  $f_0(500)\gamma$  mechanisms.

<sup>2</sup>In the model assuming the  $\rho \rightarrow \pi^0\pi^0\gamma$  decay via the  $\omega\pi$  mechanism only.

<sup>3</sup>Superseded by ACHASOV 02F.

 $\Gamma(\pi^0\pi^0\gamma)/\Gamma(\pi^+\pi^-\pi^0)$   $\Gamma_{13}/\Gamma_1$ 

<u>VALUE</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<0.00045	90	DOLINSKY 89	ND	$e^+e^- \rightarrow \pi^0\pi^0\gamma$
$\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$				
<0.08	95	JACQUET 69B	HLBC	$2.05\pi^+p \rightarrow \pi^+\rho\omega$

$\Gamma(\pi^0\pi^0\gamma)/\Gamma(\pi^0\gamma)$  $\Gamma_{13}/\Gamma_2$ 

<u>VALUE</u> (units $10^{-4}$ )	<u>CL%</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b><math>8.0 \pm 1.3</math> OUR FIT</b>					
<b><math>8.5 \pm 2.9</math></b>		$40 \pm 14$	ALDE	94B GAM2	$38\pi^- p \rightarrow \pi^0\pi^0\gamma n$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
< 50	90		DOLINSKY	89	$e^+e^- \rightarrow \pi^0\pi^0\gamma$
<1800	95		KEYNE	76	$\pi^- p \rightarrow \omega n$
<1500	90		BENAKSAS	72C OSPK	$e^+e^-$
<1400			BALDIN	71	$HLBC 2.9\pi^+ p$
<1000	90		BARMIN	64	$HLBC 1.3-2.8\pi^- p$

 $\Gamma(\pi^0\pi^0\gamma)/\Gamma(\text{neutrals})$  $\Gamma_{13}/(\Gamma_2+\Gamma_4)$ 

<u>VALUE</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$0.22 \pm 0.07$		<sup>1</sup> DAKIN	72	$OSPK 1.4\pi^- p \rightarrow nMM$
<0.19	90	DEINET	69B	OSPK

<sup>1</sup> See  $\Gamma(\pi^0\gamma)/\Gamma(\text{neutrals})$ . $\Gamma(\eta\pi^0\gamma)/\Gamma_{\text{total}}$  $\Gamma_{14}/\Gamma$ 

<u>VALUE</u> (units $10^{-5}$ )	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>&lt;3.3</b>	90	AKHMETSHIN 04B	CMD2	$0.6-0.97 e^+e^- \rightarrow \eta\pi^0\gamma$

 $\Gamma(\mu^+\mu^-)/\Gamma_{\text{total}}$  $\Gamma_{15}/\Gamma$ 

<u>VALUE</u> (units $10^{-5}$ )	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b><math>9.0 \pm 3.1</math> OUR FIT</b>				
<b><math>9.0 \pm 2.9 \pm 1.1</math></b>	18	HEISTER	02C ALEP	$Z \rightarrow \mu^+\mu^- + X$

 $\Gamma(\mu^+\mu^-)/\Gamma(\pi^+\pi^-\pi^0)$  $\Gamma_{15}/\Gamma_1$ 

<u>VALUE</u> (units $10^{-3}$ )	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>&lt;0.2</b>	90	WILSON	69	$OSPK 12\pi^- C \rightarrow Fe$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<1.7	74	FLATTE	66	$HBC 1.2-1.7 K^- p \rightarrow \Lambda\mu^+\mu^-$
<1.2		BARBARO-...	65	$HBC 2.7 K^- p$

 $\Gamma(\pi^0\mu^+\mu^-)/\Gamma(\mu^+\mu^-)$  $\Gamma_7/\Gamma_{15}$ 

<u>VALUE</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
• • • We do not use the following data for averages, fits, limits, etc. • • •				
1.2 $\pm$ 0.6	30	<sup>1</sup> DZHELYADIN	79	$CNTR 25-33\pi^- p$

<sup>1</sup> Superseded by DZHELYADIN 81B result above. $\Gamma(3\gamma)/\Gamma_{\text{total}}$  $\Gamma_{16}/\Gamma$ 

<u>VALUE</u> (units $10^{-4}$ )	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>&lt;1.9</b>	95	<sup>1</sup> ABELE	97E CBAR	$0.0\bar{p}p \rightarrow 5\gamma$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<2	90	<sup>1</sup> PROKOSHIN	95 GAM2	$38\pi^- p \rightarrow 3\gamma n$

<sup>1</sup> From direct  $3\gamma$  decay search.

### $\Gamma(\eta\pi^0)/\Gamma_{\text{total}}$

Violates  $C$  conservation.

### $\Gamma_{17}/\Gamma$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
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• • • We do not use the following data for averages, fits, limits, etc. • • •

$$<0.001 \quad 90 \quad \text{ALDE} \quad 94\text{B GAM2} \quad 38\pi^- p \rightarrow \eta\pi^0 n$$

### $[\Gamma(\eta\gamma) + \Gamma(\eta\pi^0)]/\Gamma(\pi^+\pi^-\pi^0)$

### $(\Gamma_5 + \Gamma_{17})/\Gamma_1$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<0.016	90	<sup>1</sup> FLATTE	66 HBC	$1.2 - 1.7 K^- p \rightarrow \Lambda\pi^+\pi^- \text{ MM}$

• • • We do not use the following data for averages, fits, limits, etc. • • •

$$<0.045 \quad 95 \quad \text{JACQUET} \quad 69\text{B HLBC} \quad 2.05\pi^+ p \rightarrow \pi^+ p\omega$$

<sup>1</sup> Restated by us using  $B(\eta \rightarrow \text{charged modes}) = 29.2\%$ .

### $\Gamma(\eta\pi^0)/\Gamma(\pi^0\gamma)$

Violates  $C$  conservation.

### $\Gamma_{17}/\Gamma_2$

VALUE (units $10^{-3}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<2.6	90	<sup>1</sup> STAROSTIN 09	CRYM	$\gamma p \rightarrow \eta\pi^0 p$

<sup>1</sup> STAROSTIN 09 reports  $[\Gamma(\omega(782) \rightarrow \eta\pi^0)/\Gamma(\omega(782) \rightarrow \pi^0\gamma)] \times [B(\eta \rightarrow 2\gamma)] < 1.01 \times 10^{-3}$  which we divide by our best value  $B(\eta \rightarrow 2\gamma) = 39.41 \times 10^{-2}$ .

### $\Gamma(2\pi^0)/\Gamma(\pi^0\gamma)$

### $\Gamma_{18}/\Gamma_2$

Violates  $C$  conservation and Bose-Einstein statistics.

VALUE (units $10^{-3}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<2.59	90	STAROSTIN 09	CRYM	$\gamma p \rightarrow 2\pi^0 p$

### $\Gamma(3\pi^0)/\Gamma_{\text{total}}$

Violates  $C$  conservation.

### $\Gamma_{19}/\Gamma$

• • • We do not use the following data for averages, fits, limits, etc. • • •

$$<3 \times 10^{-4} \quad 90 \quad \text{PROKOSHKIN 95} \quad \text{GAM2} \quad 38\pi^- p \rightarrow 3\pi^0 n$$

### $\Gamma(3\pi^0)/\Gamma(\pi^0\gamma)$

Violates  $C$  conservation.

### $\Gamma_{19}/\Gamma_2$

VALUE (units $10^{-3}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<2.72	90	STAROSTIN 09	CRYM	$\gamma p \rightarrow 3\pi^0 p$

### $\Gamma(3\pi^0)/\Gamma(\pi^+\pi^-\pi^0)$

Violates  $C$  conservation.

### $\Gamma_{19}/\Gamma_1$

• • • We do not use the following data for averages, fits, limits, etc. • • •

$$<0.009 \quad 90 \quad \text{BARBERIS 01} \quad 450 pp \rightarrow pf 3\pi^0 ps$$

## PARAMETER $\Lambda$ IN $\omega \rightarrow \pi^0\mu^+\mu^-$ DECAY

In the pole approximation the electromagnetic transition form factor for a resonance of mass  $M$  is given by the expression:

$$|F|^2 = (1 - M^2/\Lambda^2)^{-2},$$

where for the parameter  $\Lambda$  vector dominance predicts  $\Lambda = M_p \approx 0.770 \text{ GeV}$ . The ARNALDI 09 measurement is in obvious conflict with this expectation. Note that

for  $\eta \rightarrow \mu^+ \mu^- \gamma$  decay ARNALDI 09 and DZHELYADIN 80 obtain the value of  $\Lambda$  consistent with vector dominance.

VALUE (GeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.668±0.009±0.003</b>	3k	ARNALDI	09	NA60 158A In-In collisions
<b>• • • We do not use the following data for averages, fits, limits, etc. • • •</b>				
0.65 ± 0.03		DZHELYADIN	81B	CNTR 25–33 $\pi^- p \rightarrow \omega n$

## ω(782) REFERENCES

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BENAYOUN	13	EPJ C73 2453	M. Benayoun, P. David, L. DelBuono (PARIN, BERLIN+)	
DAIER	13	EPJ C73 2597	M. Davier <i>et al.</i>	
LEES	12G	PR D86 032013	J.P. Lees <i>et al.</i>	(BABAR Collab.)
NIECKNIG	12	EPJ C72 2014	F. Niecknig, B. Kubis, S.P. Schneider	(BONN)
BENAYOUN	10	EPJ C65 211	M. Benayoun <i>et al.</i>	
ACHASOV	09A	JETP 109 379	M.N. Achasov <i>et al.</i>	(SND Collab.)
		Translated from ZETF 136 442.		
ARNALDI	09	PL B677 260	R. Arnaldi <i>et al.</i>	(NA60 Collab.)
STAROSTIN	09	PR C79 065201	A. Starostin <i>et al.</i>	(Crystal Ball Collab. at MAMI)
ACHASOV	08	JETP 107 61	M.N. Achasov <i>et al.</i>	(SND Collab.)
		Translated from ZETF 134 80.		
AMBROSINO	08G	PL B669 223	F. Ambrosino <i>et al.</i>	(KLOE Collab.)
ACHASOV	07B	PR D76 077101	M.N. Achasov <i>et al.</i>	(SND Collab.)
AKHMETSHIN	07	PL B648 28	R. Akhmetshin <i>et al.</i>	(Novosibirsk CMD-2 Collab.)
ACHASOV	06	JETP 103 380	M.N. Achasov <i>et al.</i>	(Novosibirsk SND Collab.)
		Translated from ZETF 130 437.		
ACHASOV	06A	PR D74 014016	M.N. Achasov <i>et al.</i>	(SND Collab.)
AULCHENKO	06	JETPL 84 413	V.M. Aulchenko <i>et al.</i>	(Novosibirsk CMD-2 Collab.)
		Translated from ZETFP 84 491.		
ACHASOV	05A	JETP 101 1053	M.N. Achasov <i>et al.</i>	(Novosibirsk SND Collab.)
		Translated from ZETF 128 1201.		
AKHMETSHIN	05	PL B605 26	R.R. Akhmetshin <i>et al.</i>	(Novosibirsk CMD-2 Collab.)
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AULCHENKO	05	JETPL 82 743	V.M. Aulchenko <i>et al.</i>	(Novosibirsk CMD-2 Collab.)
		Translated from ZETFP 82 841.		
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AUBERT,B	04N	PR D70 072004	B. Aubert <i>et al.</i>	(BABAR Collab.)
PDG	04	PL B592 1	S. Eidelman <i>et al.</i>	(PDG Collab.)
ACHASOV	03	PL B559 171	M.N. Achasov <i>et al.</i>	(Novosibirsk SND Collab.)
ACHASOV	03D	PR D68 052006	M.N. Achasov <i>et al.</i>	(Novosibirsk SND Collab.)
ALOISIO	03	PL B561 55	A. Aloisio <i>et al.</i>	(KLOE Collab.)
BENAYOUN	03	EPJ C29 397	M. Benayoun <i>et al.</i>	
ACHASOV	02E	PR D66 032001	M.N. Achasov <i>et al.</i>	(Novosibirsk SND Collab.)
ACHASOV	02F	PL B537 201	M.N. Achasov <i>et al.</i>	(Novosibirsk SND Collab.)
AKHMETSHIN	02	PL B527 161	R.R. Akhmetshin <i>et al.</i>	(Novosibirsk CMD-2 Collab.)
ALOISIO	02D	PL B537 21	A. Aloisio <i>et al.</i>	(KLOE Collab.)
HEISTER	02C	PL B528 19	A. Heister <i>et al.</i>	(ALEPH Collab.)
ACHASOV	01E	PR D63 072002	M.N. Achasov <i>et al.</i>	(Novosibirsk SND Collab.)
AKHMETSHIN	01B	PL B509 217	R.R. Akhmetshin <i>et al.</i>	(Novosibirsk CMD-2 Collab.)
BARBERIS	01	PL B507 14	D. Barberis <i>et al.</i>	
ACHASOV	00	EPJ C12 25	M.N. Achasov <i>et al.</i>	(Novosibirsk SND Collab.)
ACHASOV	00D	JETPL 72 282	M.N. Achasov <i>et al.</i>	(Novosibirsk SND Collab.)
		Translated from ZETFP 72 411.		
ACHASOV	00G	JETPL 71 355	M.N. Achasov <i>et al.</i>	(Novosibirsk SND Collab.)
		Translated from ZETFP 71 519.		
AKHMETSHIN	00C	PL B476 33	R.R. Akhmetshin <i>et al.</i>	(Novosibirsk CMD-2 Collab.)
AULCHENKO	00A	JETP 90 927	V.M. Aulchenko <i>et al.</i>	(Novosibirsk SND Collab.)
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CASE	00	PR D61 032002	T. Case <i>et al.</i>	(Crystal Barrel Collab.)
ACHASOV	99E	PL B462 365	M.N. Achasov <i>et al.</i>	(Novosibirsk SND Collab.)
GARDNER	99	PR D59 076002	S. Gardner, H.B. O'Connell	
BENAYOUN	98	EPJ C2 269	M. Benayoun <i>et al.</i>	(IPNP, NOVO, ADLD+)
ABELE	97E	PL B411 361	A. Abele <i>et al.</i>	(Crystal Barrel Collab.)
BENAYOUN	96	ZPHY C72 221	M. Benayoun <i>et al.</i>	(IPNP, NOVO)
PROKOSHKIN	95	SPD 40 273	Y.D. Prokoshkin, V.D. Samoilenko	(SERP)
		Translated from DANS 342 610.		

WURZINGER	95	PR C51 443	R. Wurzinger <i>et al.</i>	(BONN, ORSAY, SACL+)
ALDE	94B	PL B340 122	D.M. Alde <i>et al.</i>	(SERP, BELG, LANL, LAPP+)
AMSLER	94C	PL B327 425	C. Amsler <i>et al.</i>	(Crystal Barrel Collab.)
ALDE	93	PAN 56 1229	D.M. Alde <i>et al.</i>	(SERP, LAPP, LANL, BELG+)
		Translated from YAF 56	137.	
Also		ZPHY C61 35	D.M. Alde <i>et al.</i>	(SERP, LAPP, LANL, BELG+)
AMSLER	93B	PL B311 362	C. Amsler <i>et al.</i>	(Crystal Barrel Collab.)
WEIDENAUER	93	ZPHY C59 387	P. Weidenauer <i>et al.</i>	(ASTERIX Collab.)
ANTONELLI	92	ZPHY C56 15	A. Antonelli <i>et al.</i>	(DM2 Collab.)
DOLINSKY	91	PRPL 202 99	S.I. Dolinsky <i>et al.</i>	(NOVO)
WEIDENAUER	90	ZPHY C47 353	P. Weidenauer <i>et al.</i>	(ASTERIX Collab.)
DOLINSKY	89	ZPHY C42 511	S.I. Dolinsky <i>et al.</i>	(NOVO)
BITYUKOV	88B	SJNP 47 800	S.I. Bityukov <i>et al.</i>	(SERP)
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DOLINSKY	88	SJNP 48 277	S.I. Dolinsky <i>et al.</i>	(NOVO)
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KURDADZE	88	JETPL 47 512	L.M. Kurdadze <i>et al.</i>	(NOVO)
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AULCHENKO	87	PL B186 432	V.M. Aulchenko <i>et al.</i>	(NOVO)
BARKOV	87	JETPL 46 164	L.M. Barkov <i>et al.</i>	(NOVO)
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KURDADZE	86	JETPL 43 643	L.M. Kurdadze <i>et al.</i>	(NOVO)
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BARKOV	85	NP B256 365	L.M. Barkov <i>et al.</i>	(NOVO)
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DZHELYADIN	81B	PL 102B 296	R.I. Dzhelyadin <i>et al.</i>	(SERP)
CORDIER	80	NP B172 13	A. Cordier <i>et al.</i>	(LALO)
DZHELYADIN	80	PL 94B 548	R.I. Dzhelyadin <i>et al.</i>	(SERP)
ROOS	80	LNC 27 321	M. Roos, A. Pellinen	(HELS)
BENKHEIRI	79	NP B150 268	P. Benkheiri <i>et al.</i>	(EPOL, CERN, CDEF+)
DZHELYADIN	79	PL 84B 143	R.I. Dzhelyadin <i>et al.</i>	(SERP)
COOPER	78B	NP B146 1	A.M. Cooper <i>et al.</i>	(TATA, CERN, CDEF+)
QUENZER	78	PL 76B 512	A. Quenzer <i>et al.</i>	(LALO)
VANAPEL...	78	NP B133 245	G.W. van Apeldoorn <i>et al.</i>	(ZEEM)
WICKLUND	78	PR D17 1197	A.B. Wicklund <i>et al.</i>	(ANL)
ANDREWS	77	PRL 38 198	D.E. Andrews <i>et al.</i>	(ROCH)
GESSAROLI	77	NP B126 382	R. Gessaroli <i>et al.</i>	(BGNA, FIRZ, GENO+)
KEYNE	76	PR D14 28	J. Keyne <i>et al.</i>	(LOIC, SHMP)
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AGUILAR-...	72B	PR D6 29	M. Aguilar-Benitez <i>et al.</i>	(BNL)
APEL	72B	PL 41B 234	W.D. Apel <i>et al.</i>	(KARLK, KARLE, PISA)
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BORENSTEIN	72	PR D5 1559	S.R. Borenstein <i>et al.</i>	(BNL, MICH)
BROWN	72	PL 42B 117	R.M. Brown <i>et al.</i>	(ILL, ILLC)
DAKIN	72	PR D6 2321	J.T. Dakin <i>et al.</i>	(PRIN)
RATCLIFF	72	PL 38B 345	B.N. Ratcliff <i>et al.</i>	(SLAC)
ALVENSLEB...	71C	PRL 27 888	H. Alvensleben <i>et al.</i>	(DESY)
BALDIN	71	SJNP 13 758	A.B. Baldin <i>et al.</i>	(ITEP)
		Translated from YAF 13	1318.	
BEHREND	71	PRL 27 61	H.J. Behrend <i>et al.</i>	(ROCH, CORN, FNAL)
BIZZARRI	71	NP B27 140	R. Bizzarri <i>et al.</i>	(CERN, CDEF)
COYNE	71	NP B32 333	D.G. Coyne <i>et al.</i>	(LRL)
MOFFEIT	71	NP B29 349	K.C. Moffeit <i>et al.</i>	(LRL, UCB, SLAC+)
ABRAMOVIC...	70	NP B20 209	M. Abramovich <i>et al.</i>	(CERN)
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DEINET	69B	PL 30B 426	W. Deinet <i>et al.</i>	(KARL, CERN)
JACQUET	69B	NC 63A 743	F. Jacquet <i>et al.</i>	(EPOL, BERG)
WILSON	69	Private Comm.	R. Wilson	(HARV)
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ASTVACAT...	68	PL 27B 45	R.G. Astvatsaturov <i>et al.</i>	(JINR, MOSU)
BOLLINI	68C	NC 56A 531	D. Bollini <i>et al.</i>	(CERN, BGNA, STRB)
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FELDMAN	67C	PR 159 1219	M. Feldman <i>et al.</i>	(PENN)
DIGIUGNO	66B	NC 44A 1272	G. Di Giugno <i>et al.</i>	(NAPL, FRAS, TRST)
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JAMES	66	PR 142 896	F.E. James, H.L. Kraybill	(YALE, BNL)
BARBARO-...	65	PRL 14 279	A. Barbaro-Galtieri, R.D. Tripp	(LRL)
BARMIN	64	JETP 18 1289	V.V. Barmin <i>et al.</i>	(ITEP)
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KRAEMER	64	PR 136 B496	R.W. Kraemer <i>et al.</i>	(JHU, NWES, WOOD)
BUSCHBECK	63	Siena Conf. 1 166	B. Buschbeck <i>et al.</i>	(VIEN, CERN, ANIK)

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